

H7(1): Vacuum Force due to Spherically Symmetric Spacetime

Consider the Lagrangian of spherically symmetric spacetime:

$$L = -mc^2 \left(m(r) - \frac{1}{c^2} \dot{\underline{r}}_1 \cdot \dot{\underline{r}}_1 \right)^{1/2} + \frac{m\hbar b}{r_1} \quad (1)$$

where

$$r_1 = \frac{r}{m(r)^{1/2}} \quad (2)$$

is given in UFT 4.16. The Euler Lagrange equation

$$\text{s:} \quad \frac{d}{dt} \frac{\partial L}{\partial \dot{\underline{r}}_1} = \frac{\partial L}{\partial \underline{r}_1} = \underline{\nabla}_1 L \quad (3)$$

where

$$\underline{\nabla}_1 L := \frac{\partial L}{\partial r_1} \underline{e}_r \quad (4)$$

By definition:

$$\underline{p}_1 := \frac{\partial L}{\partial \dot{\underline{r}}_1} \quad (5)$$

is the linear momentum.

The force equation is therefore:

$$\underline{F}_1 = \frac{d\underline{p}_1}{dt} = \frac{\partial L}{\partial r_1} \underline{e}_r \quad (6)$$

where

$$\underline{p}_1 = \gamma m \dot{\underline{r}}_1 = \frac{\gamma m \dot{\underline{r}}}{m(r)^{1/2}} \quad (7)$$

Here:

$$\gamma := \left(m(r) - \frac{1}{c^2} \dot{\underline{r}}_1 \cdot \dot{\underline{r}}_1 \right)^{-1/2} \quad (8)$$

As spherical spacetime approaches Minkowski spacetime:

$$m(r) \rightarrow 1 - (9)$$

Therefore:

$$\frac{dL}{dr_1} = -\frac{mc^2}{2} \gamma \frac{dm(r)}{dr_1} - \frac{mM_G}{r_1^2} - (10)$$

From eqs. (6) and (10):

$$\underline{F}_1 = \frac{d\underline{p}_1}{dt} = \left(-\frac{mc^2}{2} \gamma \frac{dm(r)}{dr_1} - \frac{mM_G}{r_1^2} \right) \underline{e}_r - (11)$$

$$= \underline{F} + \underline{F}(\text{vacuum})$$

where

$$\underline{F} = -\frac{mM_G}{r_1^2} \underline{e}_r - (12)$$

is the inverse square law is spherically symmetric spacetime, and

$$\underline{F}(\text{vacuum}) = -\frac{mc^2}{2} \gamma \frac{dm(r)}{dr_1} \underline{e}_r - (13)$$

the force due to spherically symmetric spacetime itself, the vacuum force.

The total force is therefore the ECE2 force

$$\underline{F} = -\underline{\nabla} \underline{\Phi} + \underline{\Omega}_r \underline{\Phi} - (14)$$

Here $\underline{\Phi}$ is the gravitational potential.

$$\underline{\Phi} = -\underline{mMG} - (15)$$

and $\underline{\Omega}_r$ is the relevant spin connection for spherically symmetric spacetime:

$$\underline{\Omega}_r = \underline{\Omega}_r \underline{e}_r - (16)$$

Therefore the vacuum force is:

$$\underline{F}(\text{vacuum}) = \underline{\Omega}_r \underline{\Phi} = -\frac{mc^2}{2} \gamma \frac{dm(r)}{dr} \underline{e}_r - (17)$$

In this equation:

$$\frac{dm(r)}{dr} = \frac{dm(r)}{dr} \frac{dr}{dr} - (18)$$

From eq. (2):

$$\begin{aligned} \frac{dr}{dr} &= \frac{d}{dr} \left(\frac{r}{m(r)} \right) - (19) \\ &= \left(m(r) - r \frac{dm(r)}{dr} \right) / m^2(r) \end{aligned}$$

So:

$$\underline{F}(\text{vac}) = -\frac{mc^2}{2} \frac{\gamma}{m^2(r)} \frac{dm(r)}{dr} \left(m(r) - \frac{dm(r)}{dr} \right) - (20)$$

The magnitude of the vacuum force due to locally symmetric spacetime.

In eq. (20):

$$\gamma = \left(n(r) - \frac{v^2}{n(r)c^2} \right)^{-1/2} \quad (21)$$

here in (r, ϕ) coordinates:

$$v^2 = \dot{r}^2 + r^2 \dot{\phi}^2 \quad (22)$$

, the Newtonian velocity.

If:

$$n(r) = 1 \quad (23)$$

then

$$F(vac) = 0 \quad (24)$$

and the orbit is an ellipse or conic section. In general the orbit is given by numerical solution of

$$\frac{dH}{dt} = 0 \quad (25)$$

$$\frac{dL}{dt} = 0 \quad (26)$$

where:

$$H = \gamma n(r) mc^2 \quad (27)$$

and

$$L = \frac{\gamma m r^2}{n(r)} \dot{\phi} \quad (28)$$

These concepts are consistent with the fact that spin connection of ERE2 describes departures from Minkowski spacetime, which is true despite vacuum energy and any non-Newtonian orbit.