

114(5) : Triple (use) Kepler with the Hamiltonian Method.

The ECE2 covariant Hamiltonian is:

$$H = \gamma mc^2 - \frac{m\Gamma G}{r} \quad - (1)$$

The Hamiltonian is a constant of motion  $\dot{H} = 0$  so:

$$\frac{dH}{dt} = 0 \quad - (2)$$

From eqs. (1) and (2)

$$mc^2 \frac{d\gamma}{dt} = m\Gamma G \frac{d}{dt} \left( \frac{1}{r} \right) \quad - (3)$$

Now use:

$$\frac{d}{dt} \left( \frac{1}{r} \right) = - \frac{\dot{r}}{r^2} \quad - (4)$$

where

$$v = \dot{r} \quad - (5)$$

So

$$c^2 \frac{d\gamma}{dt} = -v \frac{\Gamma G}{r^2} \quad - (6)$$

The relativistic force is:

$$F = \frac{d}{dt} (\gamma m v) = m v \frac{d\gamma}{dt} + m \gamma \frac{dv}{dt} \quad - (7)$$

So

$$v \frac{d\gamma}{dt} + \gamma \frac{dv}{dt} = - \frac{\Gamma G}{r^2} \quad - (8)$$

i.e.

$$- \frac{v^2}{c^2} \frac{\Gamma G}{r^2} + \gamma \frac{dv}{dt} = - \frac{\Gamma G}{r^2} \quad - (9)$$

and

$$\gamma \frac{dv}{dt} = - \frac{\Gamma G}{r^2} \left( 1 - \frac{v^2}{c^2} \right) \quad - (10)$$

so

$$\boxed{\gamma^3 \frac{dv}{dt} = - \frac{\Gamma G}{r^2}} \quad - (11)$$

which is the same result as the kinetic and Lagrangian methods.