

414(3): Relativistic Hamiltonian Method

The relativistic Hamiltonian is:

$$H = \gamma mc^2 - \frac{mMG}{r} \quad (1)$$

and is a constant of motion. So:

$$\frac{dH}{dt} = 0 \quad (2)$$

In  $\mathcal{Q}(r, \phi')$  system:

$$\gamma = \left(1 - \frac{v^2}{c^2}\right)^{-1/2} \quad (3)$$

$$v^2 = \dot{r}^2 + r^2 \dot{\phi}'^2 \quad (4)$$

where

$$\text{Here: } \phi' = \phi + \omega_1 t \quad (5)$$

It follows that:

$$\frac{d\gamma}{dt} + \frac{mG}{c^2} \frac{\dot{r}}{r^2} = 0 \quad (6)$$

where we have used:

$$\frac{d}{dt} \left( \frac{1}{r} \right) = \dot{r} \frac{d}{dr} \left( \frac{1}{r} \right) = -\frac{\dot{r}}{r^2} \quad (7)$$

Define:

$$\gamma = f^{-1/2} \quad (8)$$

then

$$\begin{aligned} \frac{d\gamma}{dt} &= \frac{d\gamma}{df} \frac{df}{dt} = -\frac{1}{2} \gamma^3 \frac{d}{dt} \left( 1 - \frac{\dot{r}^2 + r^2 \dot{\phi}'^2}{c^2} \right) \\ &= \frac{\gamma^3}{c^2} \left( \dot{r} \ddot{r} + r \dot{\phi}'^2 \dot{r} + r^2 \dot{\phi}' \ddot{\phi}' \right) \quad (8) \end{aligned}$$

From eqs. (6) and (8):

$$\gamma^3 (\ddot{r} + r \dot{\phi}'^2 + r^2 \dot{\phi}' \ddot{\phi}') = -\frac{mG}{r^2} \dot{r} \quad (9)$$

$$\text{i.e. } \gamma^3 (\ddot{r} - r \dot{\phi}'^2 + 2r \dot{\phi}' \dot{\phi}' + r^2 \dot{\phi}' \ddot{\phi}') = -\frac{mG}{r^2} \dot{r} \quad (10)$$

Possible solutions of eq. (10) are:

$$\gamma^3 (\ddot{r} - r \dot{\phi}'^2) = -\frac{mG}{r^2} \quad (11)$$

and

$$\gamma^3 (2\dot{\phi}' \dot{r} + r \ddot{\phi}') = 0 \quad (12)$$

These are the fully relativistic equations of a precessing and shrinking orbit.

They can be written as:

$$\ddot{r} - r \dot{\phi}'^2 = -\frac{mG}{\gamma^3 r^2} \quad (13)$$

and

$$2\dot{\phi}' \dot{r} + r \ddot{\phi}' = 0 \quad (14)$$

using:

$$\dot{\phi}'^2 = \left( \dot{\phi} + \omega_1 + t \frac{d\omega_1}{dt} \right)^2 \quad (15)$$

transforms eqs. (13) and (14) into the observer frame  $(r, \phi)$  giving:

$$\ddot{r} - r \dot{\phi}'^2 = -\frac{mG}{\gamma^3 r^2} + \Omega_r r \quad (16)$$

$$2\dot{r}(\dot{\phi} + \omega_1 + t\dot{\omega}_1) + r(\ddot{\phi} + 2\dot{\omega}_1 + t\ddot{\omega}_1) = 0 \quad (17)$$

where:

$$\Omega_r = \frac{-L}{mMGr} \frac{(\omega_1 + t\dot{\omega}_1)(\omega_1 + t\dot{\omega}_1 + 2\omega)}{(\omega + \omega_1 + t\dot{\omega}_1)} \quad (18)$$

the spin connection.

The calculations of Note 4/14(2) are recovered at the limit:

$$\gamma \rightarrow 1. \quad (19)$$

The usual non-relativistic Hamiltonian is recovered from eq. (1) using:

$$H_0 = H - mc^2 = (\gamma - 1)mc^2 - \frac{mMGr}{r} \quad (20)$$

here:

$$T = (\gamma - 1)mc^2 \quad (21)$$

the relativistic kinetic energy. Using:

$$T = \left( \left( 1 - \frac{v^2}{c^2} \right)^{-1/2} - 1 \right) mc^2 \quad (22)$$

$$\sim \left( 1 + \frac{v^2}{c^2} + \dots - 1 \right) mc^2$$

$$= \frac{1}{2} mv^2$$

gives the non-relativistic kinetic energy when:

$$v \ll c \quad (23)$$

Simultaneous numerical solution of eqs. (16) and (17) gives the fully relativistic, spinning and precessing orbit.