

# 413(8) : Transformation of Coordinate Proof of Orbital Precession and Shifting.

The plane polar coordinates are transformed from  $(r, \phi)$  to  $(r, \phi')$ , where:

$$\phi' = \phi + \omega_1 t \quad - (1)$$

It follows that the Hamiltonian is transformed to:

$$H = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\phi}'^2) - \frac{mMG}{r} \quad - (2)$$

and that the Lagrangian is transformed to:

$$L = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\phi}'^2) + \frac{mMG}{r} \quad - (3)$$

From eq. (2):

$$\phi'(r) = \int \frac{L}{r^2} \frac{dr}{\left( 2\mu \left( H - U - \frac{L^2}{2\mu r^2} \right) \right)^{1/2}} \quad - (4)$$

$$\text{so } \cos \phi' = \frac{L^2}{\mu k r} - 1 \quad - (5)$$
$$\left( 1 + \frac{2HL^2}{\mu k^2} \right)^{1/2}$$

where  $L = \mu r^2 \dot{\phi}' \quad - (6)$

Eq. (5) is the conic section:

$$r = \frac{d}{1 + \epsilon \cos \phi'} \quad - (7)$$

which is a precessing ellipse.

1. The less equations to Hamiltonian  $H$  and total angular momentum  $L$  are constants of motion, so:

$$\frac{dH}{dt} = 0, \quad \frac{dL}{dt} = 0. \quad - (8)$$

The half right latitude is:

$$\alpha = \frac{L^2}{\mu k} \quad - (9)$$

and the ellipticity is:

$$\epsilon = \left( 1 + \frac{2HL^2}{\mu k^2} \right)^{1/2} \quad - (10)$$

where the reduced mass is:

$$\mu = \frac{mM}{m+M} \quad - (11)$$

and

$$k := mM G \quad - (12)$$

If  $T$  is the time to complete a revolution of  $2\pi$ , then the precession is:

$$\Delta\phi = \omega_p T \quad - (13)$$

This explains the precession of the Hulse Taylor pulsar and the S2 star system.

The semi major axis of the orbit is:

$$a = \frac{d}{1-\epsilon^2} = \frac{k}{2|H|} \quad - (14)$$

The semi minor axis is:

$$b = \frac{d}{(1-\epsilon^2)^{1/2}} = \frac{L}{(2\mu|H|)^{1/2}} \quad - (15)$$

3) From eqs. (1) and (2):

$$\phi' \xrightarrow[t \rightarrow \infty]{} \infty \quad - (16)$$

So:  $H \xrightarrow[t \rightarrow \infty]{} \infty \quad - (17)$

In consequence:  $a \xrightarrow[t \rightarrow \infty]{} 0 \quad - (18)$

and the orbit shrinks to zero, Q.E.D.  
 Kepler's Law for areal velocity (Kepler's second law) is:

$$\frac{dA}{dt} = \frac{1}{2} r^2 \phi' = \frac{L}{2\mu} = \text{constant} \quad - (19)$$

so  $r \xrightarrow[t \rightarrow \infty]{} 0$  and  $\phi' \xrightarrow[t \rightarrow \infty]{} \infty$  - (20)

Q.E.D. This explains the shrinking of the orbit of the Hulse Taylor binary pulsar Q.E.D.  
 gravitation and radiation, Q.E.D.

Kepler's first law is modified to Eq. (7), to give a necessary orbit.

Kepler's third law is the direct result of eq. (19):

$$dt = \frac{2\mu}{L} dA \quad - (21)$$

So:

$$4) \quad T = \int_0^t dt = \frac{2\mu}{L} \int_0^A dA \quad - (22)$$

$$\text{so} \quad T = \frac{2\mu}{L} A \quad - (23)$$

The area of the ellipse is:

$$A = \pi ab \quad - (24)$$

$$\text{where} \quad b = (da)^{1/2} \quad - (25)$$

so from Eqs. (9), (14) and (15):

$$T^2 = \frac{4\pi^2 \mu}{L^3} a^3 \quad - (26)$$

$$\text{so} \quad \boxed{\begin{array}{c} T \xrightarrow{\quad} \infty \\ \phi \rightarrow \infty \end{array}} \quad - (27)$$

Eq. (26) is Kepler's third law.

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