

42(4). New Method of Determining the True Anomaly of Planar Orbits $\theta(t)$

Consider the equation of Lagrangian dynamics for any planar orbit:

$$\frac{d^2 r}{dt^2} + \Omega^2(r)r = 0 \quad - (1)$$

Let:

$$\Omega^2 = -\frac{L_0^2}{m^2 r^4} - \frac{F(r)}{mr} \quad - (2)$$

where L_0 is the conserved total angular momentum, m is a mass orbiting around M , and F is the force between m and M . The plane polar coordinate system is (r, θ) .

The general solution of eq. (1) is:

$$t = \frac{1}{\sqrt{2}} \int \left(-1 \int r \Omega^2 dr - A \right)^{-1/2} dr + B \quad - (3)$$

where A and B are constants of integration. Now consider the elliptical orbit:

$$r = \frac{d}{1 + \epsilon \cos \theta} \quad - (4)$$

where d is the half right latitude and ϵ the eccentricity. The area of the ellipse is πab

$$ab = d^2 (1 - \epsilon^2)^{-3/2} \quad - (5)$$

where
In a time T the area of the ellipse is

2) transcribed. So:

$$A(T) = \pi ab \quad - (6)$$

In a time t the following area is transcribed:

$$A(t) = \frac{t}{T} A(T) \quad - (7)$$

(see Marion and Thornton, 3rd ed., chapter 7). From eq. (8) of the previous note:

$$A(T) = \frac{L_0 T}{2m} \quad - (8)$$

so

$$A(t) = \frac{L_0 t}{2m} \quad - (9)$$

From eq. (7) of the previous note:

$$\theta = \frac{2A}{r^2} \quad - (10)$$

so

$$\theta(t) = \frac{L_0 t}{mr^2} \quad - (11)$$

which is consistent with:

$$\frac{d\theta}{dt} = \frac{L_0}{mr^2} \quad - (12)$$

In astronomy the quantity $\theta(t)$ is known as the true anomaly.

For elliptical orbit (4):

$$L_0^2 = dm^2 \underline{M} \underline{G}, \quad - (13)$$

and

$$F(r) = - \frac{m \underline{M} \underline{G}}{r^2} \quad - (14)$$

so:

$$\Omega^2 = \frac{\underline{M} \underline{G}}{r^3} \left(1 - \frac{d}{r} \right) \quad - (15)$$

Here:

$$r_{\min} = \frac{d}{1+\epsilon}, \quad r_{\max} = \frac{d}{1-\epsilon} \quad - (16)$$

and ϵ is found from:

$$\frac{1-\epsilon}{1+\epsilon} = \frac{r_{\min}}{r_{\max}} \quad - (17)$$

Here r_{\max} is the maximum separation of m from \underline{M} and r_{\min} is the minimum separation. So d and ϵ may be found experimentally.

The true anomaly is therefore:

$$\theta(t) = \frac{L_0}{\sqrt{2} m r^2} \left[\int \left(- \int r \Omega^2 dr - A \right)^{-1/2} dr + B \right] \quad - (18)$$

for any force law.

4) The perihelion is r_{\min} , the part at which m is closest to M . At the perihelion:

$$\theta(\text{perihelion}) = \frac{L_0}{\sqrt{2} m r_{\min}^2} \int_0^{r_{\min}} \left(- \int_0^{r_{\min}} r^2 \Omega^2 dr - A \right)^{-1/2} dr + B \quad - (19)$$

For an ellipse, from eq. (10) of the previous note:

$$\theta(\text{perihelion}) = \frac{2\pi ab}{r_{\min}^2} \quad - (20)$$

Therefore, at the perihelion of an elliptical orbit:

$$\frac{L_0}{\sqrt{2} m} \int_0^{r_{\min}} \left(- \int_0^{r_{\min}} r^2 \Omega^2 dr - A \right)^{-1/2} dr + B = 2\pi ab$$

$$= 2\pi d^2 (1 - \epsilon^2)^{-3/2} \quad - (21)$$

in which
$$\Omega^2 = \frac{MG}{r^3} \left(1 - \frac{d}{r} \right) \quad - (22)$$

For an elliptical orbit, the angle traversed in one orbit, from perihelion to perihelion, is:

$$\theta = 2\pi \left(\frac{d^2 (1 - \epsilon^2)^{-3/2}}{r_{\min}^2} \right) \quad - (23)$$

5) Kepler's

$$\theta = 2\pi (1 + \epsilon)^2 (1 - \epsilon^2)^{-3/2} \quad - (24)$$

and for the circle:

$$\epsilon = 0, \theta = 2\pi \quad - (25)$$

The general expression (19) therefore ought to give eq. (24) when:

$$L_0^2 = dm^2 \underline{M}G \quad - (26)$$

$$d = (1 + \epsilon) r_{\min} \quad - (27)$$

$$\Omega^2 = \frac{\underline{M}G}{r^3} \left(1 - \frac{d}{r} \right) \quad - (28)$$

$$F = - \frac{m \underline{M} G}{r^2} \quad - (29)$$

i.e. eq. (21) ought to be true:

$$\frac{L_0}{\sqrt{2m}} \int_0^{r_{\min}} \left(- \int_0^{r_{\min}} r \Omega^2 dr - A \right)^{-1/2} dr + B = 2\pi d^2 (1 - \epsilon^2)^{-3/2} \quad - (30)$$

where $r_{\min} = \frac{d}{1 + \epsilon} \quad - (31)$

This can be used as a simple check for
the calculation of θ in the Newtonian theory.

6) Initially it may be assumed that:

$$A = B = 0 \quad - (32)$$

so:

$$\frac{L_0}{\sqrt{2}m} \int_0^{r_{\min}} \left(- \int_0^{r_{\min}} \frac{MG}{r^2} \left(1 - \frac{d}{r} \right) dr \right)^{-1/2} = 2\pi d^2 (1 - \epsilon^2)^{-3/2} \quad - (33)$$

$$L_0 = m (dMG)^{1/2} \quad - (34)$$

$$r_{\min} = \frac{d}{1 + \epsilon} \quad - (35)$$

If this check works with the assumption (32) then the next step is to change the force law, and to recalculate eq. (18) to give θ for the new force law.

This gives the precession of the perihelion for any force law for θ change is θ .